Wild Sun! A Drama in Three Acts

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#### Wild Sun! A Drama in 3 Acts

Act 1: October Storm
Scene 1: At Earth, October 2011
Scene 2: At L1
Scene 3: On the Sun, 2 days earlier

Act 2: A Tale of Two Atmospheres

Act 2: Sun-Earth Connection

#### Prelude Your relationship with the Sun



## Act 1: October Storm Scene 1

# At Earth Late October 2011

# In the News...

Epic Geomagnetic Storm Erupts Discovery News 10/25/11

Sun Storm Paints the Night Sky *Washington Post* 10/26/11

Northern Lights Seen Across Southeast US ABC News 10/25/11

Unusual Northern Lights Set Southern Skies Afire *Roanoke Times*10/26/11

Watch Out Mars! spaceweather.com 10/22/11

> Solar Flare Illuminates the Grid's Vulnerability *New York* Times 6/9/11

# At Earth, October 24, 2011







http://spaceweather.com/aurora/gallery\_01oct11.htm

## View From Over the Pole

#### Auroral oval from NOAA-15 satellite

#### October 25, 2011







Energy deposition, erg/cm<sup>2</sup>/s

Act 1 Scene 2

# At L1 (Late October 2011)

# What's L1?

#### Quick Quiz: What's L1?

- a) A stable point 60° ahead of Earth in the same orbit
- b) A point directly behind the Moon
- c) A point where Earth's and Sun's gravitational forces cancel
- d) None of the above

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## The L1 Lagrange Point

- Point where Earth's and Sun's gravity combine to give 1-year orbital period
- Located ~1 million miles sunward of Earth
- Spacecraft at L1 orbit Sun in "lockstep" with Earth
- Sun always in view



# At L1

ACE (Advanced Composition Explorer) solar wind data, late October 2011



Day of Year: 295 = October 22, 2011

Graph constructed from http://www.srl.caltech.edu/ACE/ASC/browse/view\_browse\_data.html

Act 1 Scene 3

Back at the Sun ~2 days earlier

#### Back at the Sun: October 22, 2011



STEREO COR2 coronagraph



SOHO C2 coronagraph 2011/10/22 11:24

#### Back at the Sun: October 21-22, 2011

Coronal mass ejection viewed from STEREO ahead spacecraft



http://sohowww.nascom.nasa.gov/pickoftheweek/old/28oct2011/

# Interlude: Our Eyes on the Sun





#### Solar eclipse observations





#### 06 rovide 3D imaging





# Act 2: A Tale of Two Atmospheres



**Coronal Mass Ejections:** The Big Solar Storms Most energetic events in our Solar System  $\bullet \sim 10$  trillion nuclear bombs Eject 10 trillion tons of solar material into space ◆~mass of a mountain Speeds up to 1000 miles/second









#### What Drives CMEs? Magnetic Energy



SDO AIA 171 October 21, 2011



#### Building Up Magnetic Energy

Response of the magnetic field to the Sun's differential rotation



#### Simulating a Solar Storm



Simulation by Ben Lynch, Space Sciences Lab, UC Berkeley: http://sprg.ssl.berkeley.edu/~blynch/

## Seasons on the Sun: The Solar Cycle



## Seasons on the Sun: The Solar Cycle



# Act 3: Sun-Earth Connection

# A Short Physics Lesson: Charged Particles and Magnetic Fields

- Charged particles move easily *along* magnetic fields
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- Consequence: they trace out spiral paths in magnetic fields

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## Earth's Magnetic Field

#### Simple view: Earth in isolation



#### Earth's Magnetic Field

More complex: Interaction with the solar wind

Polar cusp

Bow Shock

Magnetosphere

Solar wind

Polar cusp

#### Auroras

- Result from high-energy solar particles penetrating the polar cusps
- Particles excite oxygen & nitrogen atoms in upper atmosphere
  - Atoms de-excite, emitting light
- Particles "mirror" back and forth between northern and southern hemispheres



#### Another Physics Lesson

#### Changing magnetic fields induce electric currents

 Basis of electric generators
 Basis of geomagnetic storms



Tue, 20 April, 2010

Orbital Blames Galaxy 15 Failure on Solar Storm

#### Simulating a CME: Sun to Earth

Aurora over Ann Arbor, MI October 29, 2003

Another Strong Magnetic Storm Pummels Earth

http://helios.astro.lsa.umich.edu/~kristin/aurora2

Los Angeles Times 10/31/03

Flare Damages Mars Odyssey Probe BBC News 11/28/03



Courtesy of Ward Manchester, University of Michigan

#### The Sun and Climate

- Connection usually overblown!
- Weak solar-cycle signal present in climate records
- Total solar luminosity variation over solar cycle: ~ 1 W/m<sup>2</sup>; ~0.1%
- Resulting temperature variation: ~several hundredths of a degree
   BUT:
  - ◆ UV variation much greater
  - Forcing change in rising cycle comparable to CO<sub>2</sub> increase





#### Longer Term Effects? (~30 years)



 $\begin{array}{c} 0.6 \\ 0.5 \\ 0.4 \\ 0.3 \\ 0.2 \\ 0.1 \\ 0 \\ -0.1 \\ 1980 \\ 1990 \\ 2000 \\ 2010 \\ 2010 \end{array}$ 

Global temperature, °C (deviation from 1961-1990 average) Climatic Research Unit, UEA





# Proxy-based millennial temperature reconstructions

#### Final Scene: Sun and Earth



