## Wild Sun! A Drama in Three Acts

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# Wild Sun! A Drama in 3 Acts 

- Act 1: October Storm
- Scene 1: At Earth, October 2011
- Scene 2: At L1
- Scene 3: On the Sun, 2 days earlier
- Act 2: A Tale of Two Atmospheres
- Act 2: Sun-Earth Connection


## Prelude

## Your relationship with the Sun



# Act 1: October Storm Scene 1 

## At Earth

Late October 2011

## In the News...

Epic Geomagnetic Storm Erupts Discovery News 10/25/11

Sun Storm Paints the Night Sky Washington Post 10/26/11

Northern Lights Seen Across Southeast US ABC News 10/25/11

Unusual Northern Lights Set Southern Skies Afire Roanoke Times10/26/11

Watch Out Mars! spaceweather.com 10/22/11

Solar Flare Illuminates the Grid's Vulnerability New York Times 6/9/11

## At Earth, October 24, 2011



## View From Over the Pole

Auroral oval from NOAA-15 satellite
October 25, 2011



## Act 1 <br> Scene 2

At L1
(Late October 2011)

## What's L1?

Quick Quiz: What's L1?
a) A stable point $60^{\circ}$ ahead of Earth in the same orbit
b) A point directly behind the Moon
c) A point where Earth's and Sun's gravitational forces cancel
d) None of the above

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## The L1 Lagrange Point

- Point where Earth's and Sun's gravity combine to give 1-year orbital period
- Located $\sim 1$ million miles sunward of Earth
- Spacecraft at L1 orbit Sun in "lockstep" with Earth
- Sun always in view


## At L1

## ACE (Advanced

Composition Explorer) solar wind data, late October 2011


Day of Year: 295 = October 22, 2011

## Act 1 Scene 3

Back at the Sun ~2 days earlier

## Back at the Sun: October 22, 2011



SOHO C2 coronagraph

## Back at the Sun: October 21-22, 2011

Coronal mass ejection viewed from STEREO ahead spacecraft



## Interlude:

Our Eyes on the Sun

Ground-based solar telescopes


Hinode (Sunrise) 2006


Solar Bunamias Ohsfreffery


06
rovide 3D imaging


# Act 2: <br> A Tale of Two Atmospheres 

## Two Atmospheres

## Earth

- Thin ( $\sim 100$ miles)


## Sun (corona)

- Extended (beyond Pluto)


## )t ( $\sim 2$ million kelvin) ffuse indy irge storms ectrical conductor agnetism dominant

Coronal Mass Ejections: The Big Solar Storms

- Most energetic events in our Solar System -~10 trillion nuclear bombs

- Eject 10 trillion tons of solar material into space - ~mass of a mountain

- Speeds up to 1000 miles/second



## What Drives CMEs? Magnetic Energy



## Building Up Magnetic Energy

Response of the magnetic field to the Sun's differential rotation


## Simulating a Solar Storm



Simulation by Ben Lynch, Space Sciences Lab, UC Berkeley: http://sprg.ssl.berkeley.edu/~blynch/

## Seasons on the Sun: The Solar Cycle



## Seasons on the Sun: The Solar Cycle



## Act 3: <br> Sun-Earth Connection

## A Short Physics Lesson: <br> Charged Particles and Magnetic Fields

- Charged particles move easily along magnetic fields
- It's difficult for them to move perpendicular to magnetic fields
- Consequence: they trace out spiral paths in
 magnetic fields


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## Earth's Magnetic Field

Simple view: Earth in isolation


## Earth's Magnetic Field

More complex: Interaction with the solar wind


## Auroras

- Result from high-energy solar particles penetrating the polar cusps
- Particles excite oxygen \& nitrogen atoms in upper atmosphere
- Atoms de-excite, emitting light
- Particles "mirror" back and forth between northern and southern hemispheres



## Another Physics Lesson

- Changing magnetic fields induce electric currents
$\Delta$ Basis of electric generators
- Basis of geomagnetic storms

Salar activity domaand Marg

Tue, 20 April, 2010
Orbital Blames Galaxy 15 Failure on Solar Storm

## Simulating a CME: Sun to Earth

## Aurora over Ann

Arbor, MI
October 29, 2003

## 原

http://helios.astro.lsa.umich.edu/~kristin/aurora2
Another Strong Magnetic Storm Pummels Earth
Los Angeles Times 10/31/03

Flare Damages Mars
Odyssey Probe
BBC News 11/28/03

University of Mithigan
Manchester et, at.
2885


Courtesy of Ward Manchester, University of Michigan

## The Sun and Climate

- Connection usually overblown!
- Weak solar-cycle signal present in

climate records
- Total solar luminosity variation over solar cycle: $\sim 1 \mathrm{~W} / \mathrm{m}^{2} ; \sim 0.1 \%$
- Resulting temperature variation:


Zurich sunspot numbers ~several hundredths of a degree

- BUT:
- UV variation much greater
- Forcing change in rising cycle comparable to $\mathrm{CO}_{2}$ increase


11-year Fourier component Wolfson \& Hand SVECSE 2008

## Longer Term Effects? ( $\sim 30$ years)




Global temperature, ${ }^{\circ} \mathrm{C}$ (deviation from 1961-1990 average)

Climatic Research Unit, UEA

## Millennial effects?





Proxy-based millennial temperature reconstructions

## Final Scene: Sun and Earth

Not this...



