## Mission Design:

 Exploring the Solar SystemKathleen Howell
Purdue University

## Potential Destinations?



## Planets: Saturn?



## Dwarf Planet?

Namaka


Haumea

Makemake


## Dwarf Planet:

(a) in orbit around the Sun
(b) sufficient mass for selfgravity to assume nearly round shape
(c) neighborhood around orbit not cleared
(d) is not a satellite

## Earth



Eris


Ceres

Pluto


## Izion

| $2003 \mathrm{EL}_{61}$ | $2005 \mathrm{FY}_{8}$ | Seina | Orcus | Quanar | $2002 \mathrm{TX}_{300}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | C | - |  | 6 | e |
| $2002 \mathrm{AWH}_{10}$ | Varuna | Izion | Vesta | Pallas | Hyylea |

## Kuiper Belt Object?



Comets

## Oort Cloud Object?



The Oort Cloud
(comprising many
billions of comets)
Oort Cloud cutaway
drawing adaplod fiom ilvsitraton (NASA, JPL)

## Mission Design

Mission Architectures
Mission Requirements
Mission Planning
Technology Utilization
Cost/Performance Analysis

## Conics from Kepler and Newton



Conic Orbits: Three-dimensional Characteristics


## Dwarf Planet Eris



Orbit of Eris (136199 Eris)

Perihelion: 37.77 AU
Aphelion: 97.56 AU
Orbital period: 557 years

## Problem:

Design spacecraft trajectory $\quad \Rightarrow \quad$ specific requirements

| Approach: |
| :--- |
| Traditional Two-Body |
| - Analytical Solns $\left\{\begin{array}{l}\text { ellipses } \\ \text { parabolas } \\ \text { hyperbolas }\end{array}\right.$ |
| - Identify various trajectory arcs; <br> patch together |
| • Transition to full model |
| • Optimize in full model |



## Sputnik

Launch: October 4, 1957 $1^{\text {st }}$ in Sputnik Program

Technology First !!

Orbit: $7310 \times 6586$ km 96.2 minutes/rev


Science measurements:
Density of upper layer of atmosphere
$\rightarrow$ changes in orbit
Radio signal distribution data in ionosphere Meteoroid detection


Play \#1

ISS Trajectory




## PROJECT APOLLO



LUNAR LANDING FLIGHT TECHNIQUES


## PURDUE

U N I VERSITY


## Lunar Prospector

## Tirajectory

Orbit

## Type: Orbiter



Central Body: Earth
Epoch start: 1998-01-07 03:30:00 UTC
Orbital Parameters

| Periapsis | Apoapsis | Period | Inclination | Eccentricity |
| :--- | :--- | :--- | :--- | :--- |
| 1.03093004226 | 56.2900009155 | 216.100006103 | 29.2000007629 | 0.96403002738 |
| 68457 RE | 27344 RE | 51562 hours | $39453^{\circ}$ | 95264 |

## Trajectory Design

Combine arcs of 3 shapes: ellipses
parabolas
hyperbolas


## Trajectory Design

Combine arcs of 3 shapes: $\left.\begin{array}{l}\text { ellipses } \\ \text { parabolas } \\ \text { hyperkolas }\end{array}\right\} \begin{gathered}\text { Different } \\ \text { Energy Levels } \\ \text { 'Stable' }\end{gathered}$
Maneuver ( $\Delta V$ )

Maneuver ( $\Delta V$ )

## Hohmann Transfer



## Construction:

1. Circular \#1
2. Maneuver to Ellipse \#2
3. Remain in Ellipse \#2; return
4. Maneuver to Ellipse \#3


## Hohmann Transfer to Uranus

TOF = 16 years
$\Delta V=16 \mathrm{~km} / \mathrm{s}$

Hohmann
$\rightarrow$ not the way to get to Uranus!!! Both Voyager launches on Titan III-Centaurs; only enough energy to reach Jupiter

Both used gravity assists to reach final destinations and even out of the solar system!

Hohmann Schematic Transfer:


Earth Orbit

## Cassini to Saturn



Transfer Trajectory:
Depart Earth 10/97
Pass Venus 4/98
Pass Venus 6/99
Pass Earth 8/99
Pass Jupiter 12/00
Arrive Saturn 7/04


## Cassini to Saturn





## Venus Flyby




Cassini
Spacecraft
First Venus Swingby
April 26, 1998
$\qquad$

Jupiter Swingby December 30, 2000

Launch to 1st Venus Swingby
1st Venus Swingby to 2nd Venus Swingby
2nd Venus Swingby to Earth Swingby, Past Jupiter to Saturn


Cassini Spacecraft (Purple) Enroute to Saturn (Gold Orbit)

Cassini Spacecraft (Purple): Enroute to Saturn (Gold)


Mercury Messenger


Science and exploration goals cannot always be met using conics, even with gravity assists!!

## Innovation in Trajectory Design 1970's




## Poincaré $\rightarrow$ Three-Body Problem

Modern Computers +


Advances in Mathematics


Two assumptions expand options:

1. New perspective: View from Earth?
2. Multiple Gravity Fields

Orbit propagated for 4 conic periods:
$4 * 19$ days $=75.7$ days


## Inertial View

Orbits propagated for 4 conic periods:
4*19 days $=75.7$ days


## Inertial View

Orbit propagated for 4 conic periods: 4*19 days $=75.7$ days


Inertial Frame


Play \#2

Orbit propagated for 4 conic periods:


To Sun

Rotating View

Rotating Frame


Inertial Frame


Rotating Frame


Orbit propagated for 4 conic periods: $4 * 19$ days $=75.7$ days


## Inertial View

Orbits propagated for 4 conic periods: $4 * 19$ days $=75.7$ days


## Inertial View

Orbits propagated for 4 conic periods: $4 * 47$ days $=188$ days


Inertial View

Orbits propagated for 4 conic periods:
4*47 days $=188$ days


Inertial View

Resonant orbit propagated for 5.2 yea


Resonant orbit propagated for 5.2 years


Inertial View



## Problem:

Design spacecraft trajectory $\Rightarrow$ specific requirements

| Approaches: |  |
| :--- | :--- |
| Traditional Two-Body | N-Body Regimes (even N = 3) |
| • Analytical Solns $\left\{\begin{array}{l}\text { ellipses } \\ \text { parabolas } \\ \text { hyperbolas }\end{array}\right.$ | • No analytical solutions <br> • Limited knowledge of solution arcs |
| • Identify various trajectory arcs; <br> patch together | • Little understanding of arc <br> "overlap" |
| • Transition to full model | • Transition $\rightarrow$ propagate single <br> state |
| • Optimize in full model | • Optimizing relies on GOOD guess; |

## Poincaré $\rightarrow$ Three-Body Problem

New View + Additional
Grav Fields


## What new options exist?

1. How do we find new solutions?

Earth
2. What do they look like?
3. How do we start?

Equilibrium Solutions

Earth-Moon Distance: 384,000 km Earth Scale: 5x

Moon Scale: 10x


Play \#4

Earth-Moon Distance: 384,000 km
Earth Scale: 5x
Moon Scale: 10x

L3

fale: 15x
Scale: 100x arth Distance: 1AU


## $L_{1}$ and $L_{2}$ Lyapunov Families



## Sun-Earth System

Lyapunov Orbits

## $L_{1}$ Halo Family



## $L_{1}$ and $L_{2}$ Lyapunov Families



## $L_{1}$ and $L_{2}$ Halo Families

\&


## $L_{1}$ and $L_{2}$ Halo Families

## $\mathrm{L}_{1}, \mathrm{~L}_{2}, \mathrm{~L}_{3}$ Halo Families



## Innovation in Trajectory Design 1970's





## Relict-2

MAP



## SOHO


$\cdot L_{1}$ Halo Orbit: $A z=120,000 \mathrm{~km}, ~ A y=666,672 \mathrm{~km}$


RLP Coordinate System


Fight Dynami
NASA-GSFC


SOHO orbit schematic


## Relict-2

MAP



## PURDUE <br> UNIVERSITY

## Genesis Trajectory





## THEMIS Background: Substorms

PRIME MISSION (FY08-09) SCIENCE GOALS:

## Primary:

"How do substorms operate?"

- One of oldest, most important questions in Geophysics
- A turning point in understanding
of the dynamic magnetosphere


## First bonus science:


"What accelerates storm-time 'killer' electrons?"

- A significant contribution to space weather science


## Second bonus science:

"What controls efficiency of solar wind - magnetosphere coupling?"

- Provides global context of solar wind \& magnetosphere interaction

RESOLVING THE PHYSICS OF ONSET AND EVOLUTION OF SUBSTORMS

Principal Investigator
Vassilis Angelopoulos, UCLA
Mission Operations Manager
Manfred Bester, UCB

EPO Lead
Laura Peticolas, UCB

## THEMIS



## Primary Objective:

identify physical mechanism that leads to explosive release of energy in substorms

- 2-year mission (launch 2/07); 5 identical probes
- First NASA launch of five satellites to study substorms
- THEMIS probes align over North Am @ 4 day intervals
- Alignments - in situ measurements of particles/fields $\rightarrow$ identify region where substorm energy release; insight into process
- Successful result -- Explosion of magnetic energy at $1 / 3$ distance to moon powers substorms due to magnetic reconnection (stressed magnetic field lines suddenly "snap" to a new shape)


Blue - Earth's magnetic field over the night side White flash - energy released during substorms
$\rightarrow$ night side magnetic field acts as slingshot; propels electrons toward Earth.


## Artemis P1 /P2 Baseline Trajectory

# Trajectory Baseline <br> Lunar Gravity + Solar Perturbation + Libration Point Orbits + Lunar Orbits 



## Artemis P1 /P2 Baseline Trajectory



## P1: Phase 1



P1: Phase 1


## P1 Backflip Family



Earth-Moon Rotating Frame
(Moon-Centered)

## P1 Backflip Family

Earth-Moon Rotating Frame
(Moon-Centered)

## Backflip Stable Manifold Earth-Moon Rotating Frame




## P1 Phase 2



- DSM on 2010 March 15
- Sun-Earth $\mathrm{L}_{1}$ Lissajous Stable Manifold


## P1: Phase 2

- Max Range (6-Jun-2010)
- Trajectory and Sun-Earth $L_{1}$ Lissajous Unstable Manifold

Max Range

P1: Phases 3 and 4

Earth-Moon Rotating Frame
(Moon-Centered)

## P1 Phase 3 Two viewpoints on the $L_{1}$ to $L_{2}$ transfer

- Simultaneously matches the stable manifold surface associated with the


P1: Phases 3 and 4

Earth-Moon Rotating Frame
(Moon-Centered)

P1: Phase 4


Earth-Moọn Rotating Frame
(Moon-Centered)

## Artemis P1 /P2 Baseline Trajectory Design

